

ERRATUM: “CLASSICAL T TAURI-LIKE OUTFLOW ACTIVITY IN THE  
BROWN DWARF MASS REGIME” (2009, ApJ, 706, 1054)\*

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Figures 3–11 have an error in that the direction of the offsets should be reversed. Thus, in the spectro-astrometric plots, positive offsets are toward the top of the figure and negative offsets toward the bottom. In the position–velocity diagram shown in Figure 8, positive offsets are toward the left and negative toward the right. The result of this correction is that the orientations of the blueshifted and redshifted lobes of the outflows are now reversed. Thus, for ISO-Cha1 217 the blueshifted flow is the in northeast quadrant with a position angle (P.A.) of 20° and the redshifted flow is in the southwest quadrant (see Figure 2). Similarly, for ISO-Oph 32, the blueshifted flow now lies in the northeast quadrant with a P.A. of 60°. Table 3 has been updated accordingly.

**Table 3**  
Position Angles of the Two Outflows Observed

Object	Line	Outflow P.A. (°)
ISO-Cha1 217	[O I]λ6300	29 (±8)
	[O I]λ6363	26 (±15)
	[S II]λ6716	13 (±5)
	[S II]λ6731	20 (±6)
ISO-Oph 32	[O I]λ6300	60 (±7)

**Notes.** The P.A.s are estimated from orthogonal spectra as described in Section 3. For ISO-Cha1 217 a bipolar outflow is uncovered and estimates given above are for the blueshifted lobe. For ISO-Oph 32 only the blueshifted outflow is detected.

\* Based on data collected by UVES observations (079.C-0375(A)) at the Very Large Telescope on Cerro Paranal (Chile) which is operated by the European Southern Observatory (ESO).